

Calibration of the Calitoo photometer in laboratory

Introduction:

The aim of this document is to present the method to calibrate the sun photometer Calitoo in laboratory. Currently, we know how to calibrate that photometer by different methods but only with the sun. In it we will present the Calitoo and the different methods we know in order to finish by the method in laboratory.

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I/ Introduction of the subject

1) Presentation of the traineeship and the Calitoo

The Calitoo photometer is a manual photometer that we can take everywhere with you. Thanks to it, you can determinate the AOD. It uses 3 wavelengths (465nm, 540nm and 619nm) and it gives us some information like GPS position, time, pressure, the AOD and the Angström exponent. In order to calculate the AOD values, it uses some C_n which are some values which is proportional to the intensity of the source of light. These C_n go from 0 to 4096 and the precision announced by the web PHOTONS/LOA for the AOD is 0.02.



The current methods of calibration are the absolute calibration with Langley method or intercalibration with a photometer master of AERONET. These methods are a very long task and can't be done at every moment because we need no cloud, so we try to found another method in laboratory to calibrate these photometers.

The society TENUM which produces the Calitoo develop also a soft for the Calitoo:



On this soft we can study the data of different Calitoo but also calibrate one of them. We can find on it different tools like intercalibration, the AOD calculator, data visualization, Langley calibration,... Here we will only use the data visualization and the intercalibration:



2) AOD calculation

Before knowing how to calculate the AOD, it is important to know what it is. The optical depth is a way to measure the transparency of the atmosphere. The aerosols can alter the transparency of this atmosphere and so the AOD describe how aerosol have an impact on the transparency for different wavelengths.

We know where we can see the values of the AOD with the Calitoo but it is important to understand how it works. From the Cn data, we can find again the AOD values. These calculations are done on an excel program which follows. In order to complete this folder, we need some coefficients like the R^2 and the Angstrom exponent (α). Their formula are:

$$\mathbf{R} = \frac{(1-0.0167)^2}{1+0.0167\cos(2\pi(\frac{D}{365}))}$$

where D is the number of the day and $\alpha = -\frac{\ln(\frac{\delta\lambda}{\lambda_0})}{\ln(\frac{\lambda}{\lambda_0})}$. We need

also the solar elevation and a coefficient m which is $m=1/\sin$ (solar elevation). After that the calculation of the AOD is $\mathbf{AOD} = -\frac{1}{m} \cdot \ln(T_\lambda \cdot R^2) - \tau_{r0\lambda} \left(\frac{p}{p_0}\right) - \tau_{0\lambda}$ where m is the coefficient linked with the solar elevation, T_λ is the transmission, $\tau_{r0\lambda}$ is the ozone optical depth and $\tau_{0\lambda}$ the Rayleigh optical depth.

With all these equations, we are able to calculate directly the AOD by hand. In fact it is what is done by the soft of calitoo. So we can compare if we have the good equation comparing our calculated AOD and the one given by the soft. Doing that we find:

Date	Time	Elevation solaire	m	d : jour de l'a	R ²	Pression	Raw 465	Raw540	Raw619	AOT 1 (465nm) C	AOT par le calcul	Erreur AOT 4	Erreur pourc
19/06/2017	08:48:02	0,802851456	1,39016359	170	1,03238754	1016	2332	2368	2013	0,1203	0,120274638	2,5362E-05	0,02108449
19/06/2017	08:49:15	0,802851456	1,39016359	170	1,03238754	1016	2317	2337	2004	0,125	0,124916554	8,3446E-05	0,06677882
19/06/2017	09:09:22	0,855211333	1,32501299	170	1,03238754	1016	2391	2405	2048	0,117	0,116940979	5,9021E-05	0,05045799
19/06/2017	09:09:48	0,855211333	1,32501299	170	1,03238754	1016	2396	2405	2054	0,1153	0,115316167	1,6167E-05	0,01402043
Data from Calitoo #45													
Date	Time	Elevation solaire	m	d : jour de l'a	R ²	Pression	Raw 465	Raw540	Raw619	AOT 1 (465nm) C	AOT par le calcul	Erreur AOT 4	Erreur pourc
19/06/2017	07:13:56	0,541052068	1,94160403	170	1,03238754	1016	1936	2191	1855	0,1366	0,136567671	3,2329E-05	0,02366969
19/06/2017	07:14:55	0,541052068	1,94160403	170	1,03238754	1016	1930	2179	1852	0,1382	0,138166342	3,3658E-05	0,02435754
19/06/2017	07:15:23	0,541052068	1,94160403	170	1,03238754	1016	1949	2199	1859	0,1333	0,133120814	0,00017919	0,13451358
19/06/2017	07:15:45	0,541052068	1,94160403	170	1,03238754	1016	1956	2208	1863	0,1312	0,131274325	7,4325E-05	0,0566338
19/06/2017	07:42:43	0,628318531	1,70130162	170	1,03238754	1016	2078	2305	1934	0,1419	0,141856379	4,3621E-05	0,0307451
19/06/2017	07:43:00	0,628318531	1,70130162	170	1,03238754	1016	2074	2293	1934	0,143	0,142988914	1,1086E-05	0,00775277
19/06/2017	07:43:10	0,628318531	1,70130162	170	1,03238754	1016	2083	2306	1932	0,14	0,140443773	0,00044377	0,31647912
19/06/2017	07:43:33	0,628318531	1,70130162	170	1,03238754	1016	2086	2311	1935	0,14	0,139597836	0,00040216	0,28767333
Data from Calitoo #62													

Here are the results for the blue channel of the Calitoo, the error is only due to the rounded error done by the soft.

3) Comparison between AOD of Calitoo and AOD of aeronet

In order to know if our Calitoo are "good", we can check if there values are similar to those of AERONET. To do that it is important to see that the wavelengths are not the same between aeronet and Calitoo. So we have to calculate the aeronet AOD to the wavelengths

of Calitoo. To do that we use: $\tau_{\lambda 0} = \tau_{\lambda} \left(\frac{\lambda}{\lambda_0}\right)^{\alpha}$ where $\tau_{\lambda 0}$ is the value of the AOD to the wavelength we want, τ_{λ} the value of the AOD we know, λ and λ_0 the linked wavelength and α , the Angstrom exponent.

We have another problem of interpolation which is the time. Sometimes, the values of aeronet and those of Calitoo are not taken in the same time. So we have to do a linear interpolation.

After that, we are able to compare the different values and see if the Calitoo is close to aeronet values. So we put that values in Excel and with a little program on it we can find:

AOT 1 (465nm)	AOT aeronet	Erreur AOT 4	Erreur pourc	AOT 2 (540nm)	AOT aeronet	Erreur AOT 5	Erreur pourc	AOT 3 (619nm)	AOT aeronet	Erreur AOT 6	Erreur pourc
0,1203	0,11270985	0,00759015	6,51487869	0,091	0,08481262	0,00618738	7,03860221	0,0766	0,07268134	0,00391866	5,25003719
0,1173	0,11270985	0,00459015	3,99126774	0,0922	0,08481262	0,00738738	8,34672178	0,0711	0,07268134	0,00158134	2,19964172
0,1075	0,10677146	0,00072854	0,6800194	0,0806	0,0801736	0,0004264	0,53043347	0,0724	0,06978985	0,00261015	3,67135572
0,1063	0,10677146	0,00047146	0,44253347	0,0808	0,0801736	0,0006264	0,77826238	0,0704	0,06978985	0,00061015	0,87045906
Calitoo #45		Data from the 19/06/2017									

AOT 1 (465nm)	AOT aeronet	Erreur AOT 4	Erreur pourc	AOT 2 (540nm)	AOT aeronet	Erreur AOT 5	Erreur pourc	AOT 3 (619nm)	AOT aeronet	Erreur AOT 6	Erreur pourc
0,113	0,11586387	0,00286387	2,50268534	0,0959	0,0885793	0,0073207	7,93660323	0,0788	0,07315253	0,00564747	7,43319754
0,1142	0,11586387	0,00166387	1,44644291	0,0899	0,0885793	0,0013207	1,47994215	0,0705	0,07315253	0,00265253	3,69298601
0,1198	0,11586387	0,00393613	3,34046004	0,0856	0,0885793	0,0029793	3,42096298	0,079	0,07315253	0,00584747	7,6863209
0,1185	0,11586387	0,00263613	2,24960332	0,0839	0,0885793	0,0046793	5,42593185	0,0781	0,07315253	0,00494747	6,54199421
0,1133	0,11365964	0,00035964	0,31692104	0,0882	0,08662729	0,00157271	1,79916208	0,0751	0,07200054	0,00309946	4,21406943
0,1158	0,11365964	0,00214036	1,8655643	0,0816	0,08662729	0,00502729	5,97677935	0,0753	0,07200054	0,00329946	4,47990136
0,1135	0,11365964	0,00015964	0,14055439	0,0884	0,08662729	0,00177271	2,02564201	0,0701	0,07200054	0,00190054	2,67492373
0,1131	0,11365964	0,00055964	0,49359879	0,0875	0,08662729	0,00087271	1,00238526	0,0754	0,07200054	0,00339946	4,6125468
Calitoo #62		Data from the 19/06/2017									

We can see here the errors between the AOD given by Calitoo and the one recalculated of AERONET. We can see that the values of the Calitoo are close to those of AERONET. So this Calitoo is considered as "good" but it's possible to find other Calitoo which values could be far from that. The values given by AERONET are with a potential error of maximum 0.01 on the AOD and the one given by Calitoo 0.02. So if we take a look of the difference between the two AOD the maximum of difference must be under

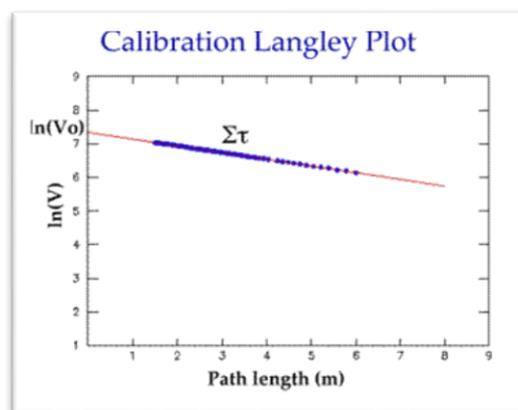
$\sqrt{0.01^2 + 0.02^2} \approx 0.022$. On these data, the maximum of error between AERONET and Calitoo is 0.007 so it is far under the 0.022 announced.

II/ Actual methods of calibration with the sun

1) Absolute calibration

The absolute calibration follows in fact the Langley calibration and doing that method, we have to take measures all the daylong in order to have a lot of measures in function of the air mass. We need to have a perfect sky and a constant AOD or zero. That is a difficulty of that method.

With all these measures and the equation of the AOD we can found the V_0 . In fact, we have the values of some V of the day and the equation: $\ln(V) = -m\text{AOD} + \ln(V_0)$ so if the m value is 0, we can find the value of V_0 and so calibrate the instrument. On this equation, the AOD is a constant that is the reason why we can do that, we have one less parameter. With all the values taken, we can draw a graph like the one which follows and we must find a straight line and the ordered at the origin gives us the value of the new V_0 .

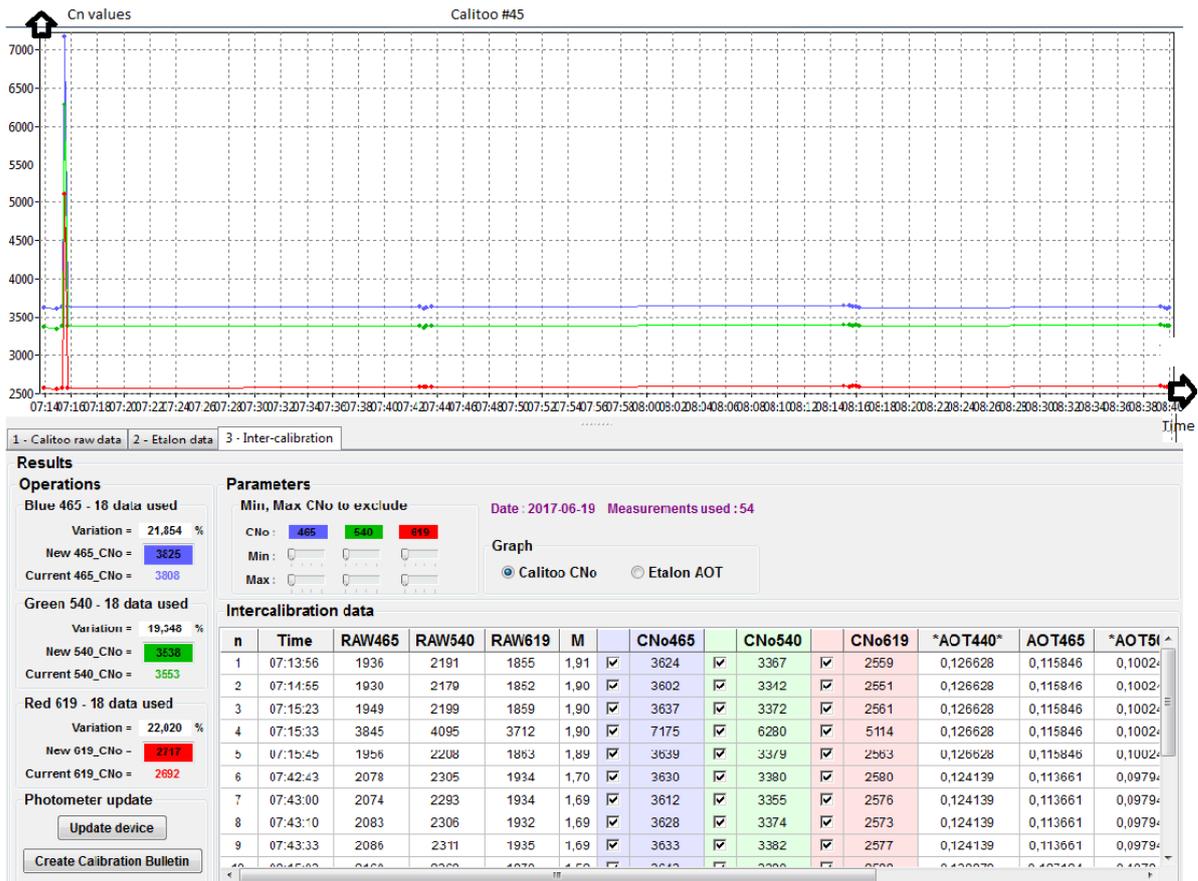


For the Calitoo this method is very repetitive because we have to take all measures by hand and by some perfect conditions. So this method was forget by the constructor and we can't do it here because it is complicated to put in place.

2) Intercalibration with a Lille photometer

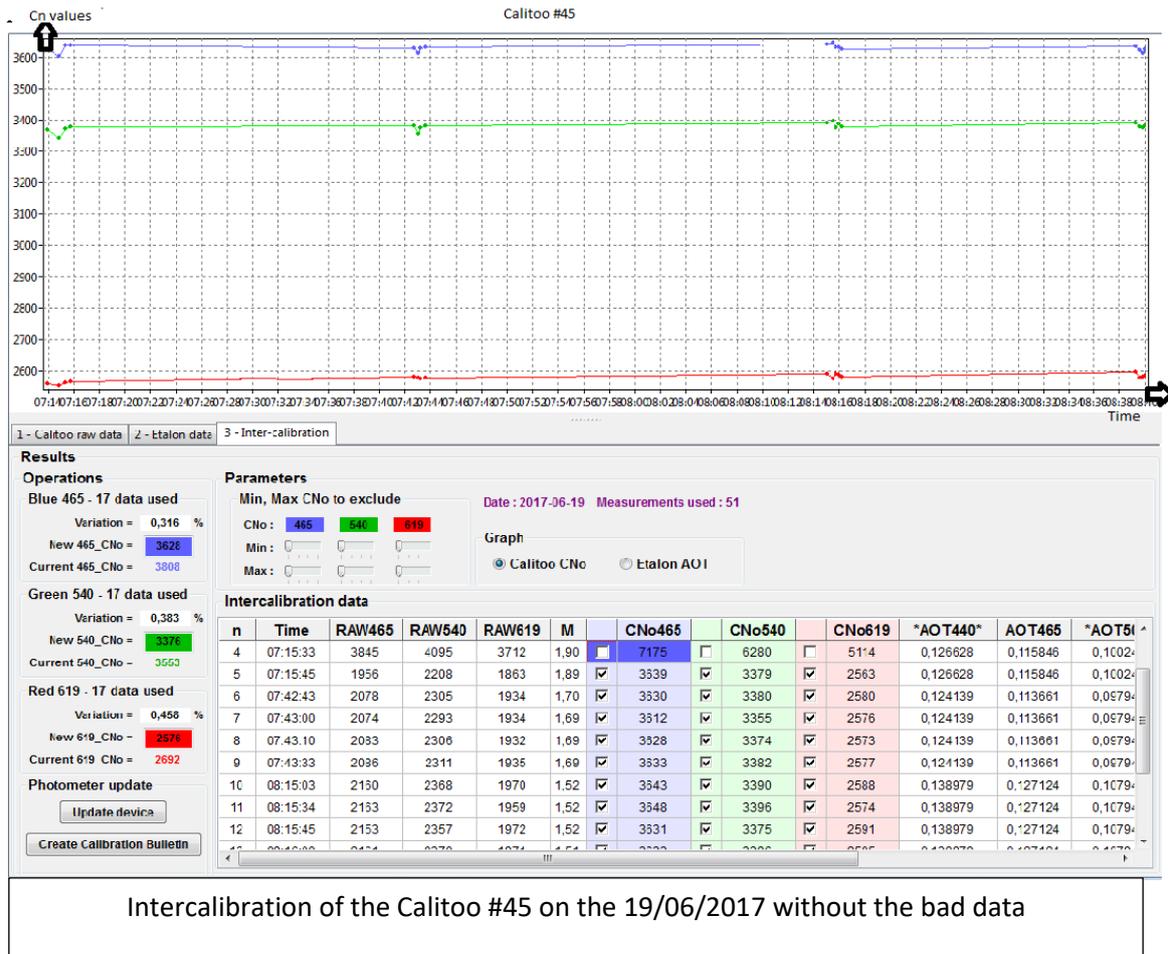
In order to calibrate our Calitoo we can use the method said as intercalibration and to begin we can intercalibrate a photometer on the roof of the PHOTONS network and a Calitoo. To put this method in place, we have to take some measures with our Calitoo on the roof near the AERONET photometer.

Our measures taken, we can analyze them. We have a tool on the soft of Calitoo which permit to intercalibrate the both. With that tool, we have only to connect the Calitoo, to choose the day when we have done the measures and to collect the AERONET data. After that the soft calculate automatically the new Cn_0 s and we can see the differences between the old and the new one or to select only the values we need.



Intercalibration of the Calitoo #45 on the 19/06/2017 without checking

For instance, on this picture we can see that the intercalibration is not very good because of one value so we can delete it to calibrate the calitoo.



Without the bad data we have a better calibration lose to old values of Cn0s so we can update the Calitoo and/or create a calibration bulletin like the one provided when you buy a Calitoo.

We can also do these calculations with a spreadsheet to understand better how it works. To be able to do that, we need also to do measures with the Calitoo but this time we will analyze them by hand. We need to know all the data of the Calitoo to calculate a coefficient which is the R^2 . Its formula is: $R = \frac{(1-0.0167)^2}{1+0.0167 \cos(2\pi(\frac{D}{365}))}$ where D is the number of the day. We can also calculate the Angstrom exponent thanks to these data. This coefficient is:

$$\alpha = - \frac{\ln(\frac{\delta\lambda}{\lambda_0})}{\ln(\frac{\lambda}{\lambda_0})}$$

We can calculate the new Cn0 as done by the soft of Calitoo and compare both. The equation of the Cn0 is: $Cn0_\lambda = Cn_\lambda \cdot R^2 \cdot \exp(m \cdot AOD_\lambda)$

Here we can see that we need some data from the Calitoo to do that in particular the old Cn0s. After that, we put the data taken at the same time between the Calitoo and we produce our ratios ($Cn_{465_1stCalitoo}/Cn_{465_2^{nd}Calitoo}$ for instance) and so we are able to calculate the new value of the Cn0 for the second Calitoo. We can see that on the previous screen and the errors between the new and the old values for that manipulation. Here the difference could appear as big between those Cn0 but in percentage it is only few percent (between 1 and 7% of difference). We are also able to calculate by hand the values of the AOD with these ratios and the results are on the screen with the associated error. This error is also about few percent (maximum of 5% of difference between old and new Cn0).

Here the Calitoo number #71 is the master and it helps us to calibrate the #45, a field instrument. The results here show that if we use the new Cn0 and these ratios, the AOD value is closer to the one of a calibrated photometer.

To conclude we can compare these two methods. About the facility of the methods we can say that the absolute calibration is much more difficult to put in place than the intercalibration. Even if the absolute calibration is the more precise method, the results of the intercalibration show that this method is reliable. The results given by the intercalibration are used nowadays to calibrate the Calitoo. So the better method between both is the intercalibration because it needs less conditions to be put in place even if we need to have a clear sky and go to Izana it takes less time and it is an easier method. In order to complete that we try to find an alternative method in laboratory.

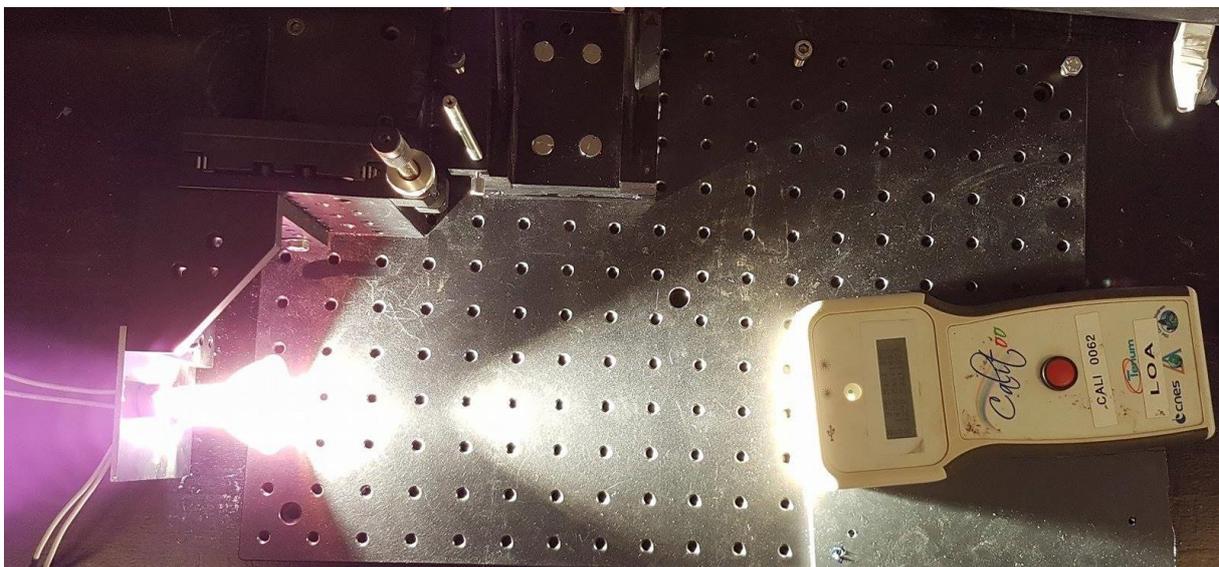
III/ New method of calibration

1) Choice of the method and material used

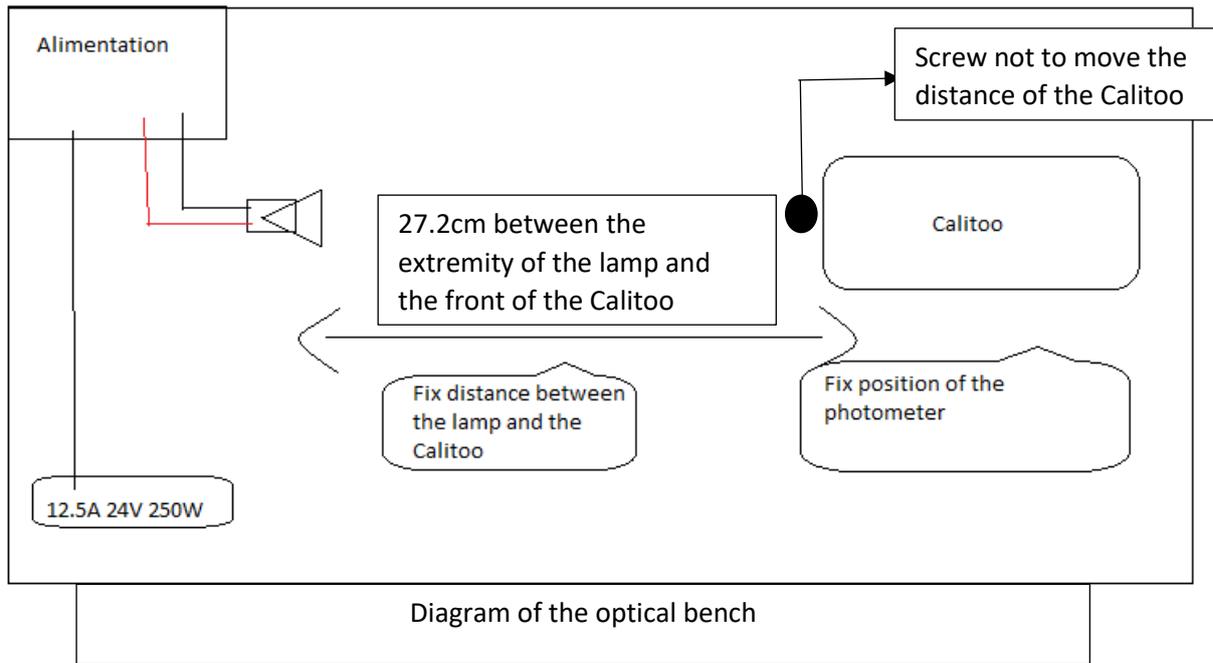
The calibration with the sun has some inconvenient that we would want to delete. In fact we can't do this calibration everywhere and every time and often, somebody has to go to the Izana (Canaries) in order to do the calibration. This is why we try to find a lab method of calibration easier to put in place. We need to calibrate the photometers regularly so it could be a gain of time and money if it is not necessary to go to Izana in order to calibrate the Calitoo.

To do that calibration, we place the calitoo in front of a lamp and we do measures. That calitoo is supposed to be calibrated and it will be our master. After that we place the other calitoo that we have to calibrate in front of the lamp and we also do measures. To finish we do again measures with the reference. All these data are analyzed after.

For the manipulation in laboratory, we need in addition to the calitoo, a powerful lamp in order to simulate the sun. This lamp is a 250W one with 12.5A and 24V. It is important to supply in consequence the lamp. Firstly, we used some batteries to know if the lamp was powerful enough for our Calitoo. After that we decided to buy a new supply which is more stable and less dangerous than the batteries. On this manipulation, we place the Calitoo in front of the lamp as we can see on the following photo:



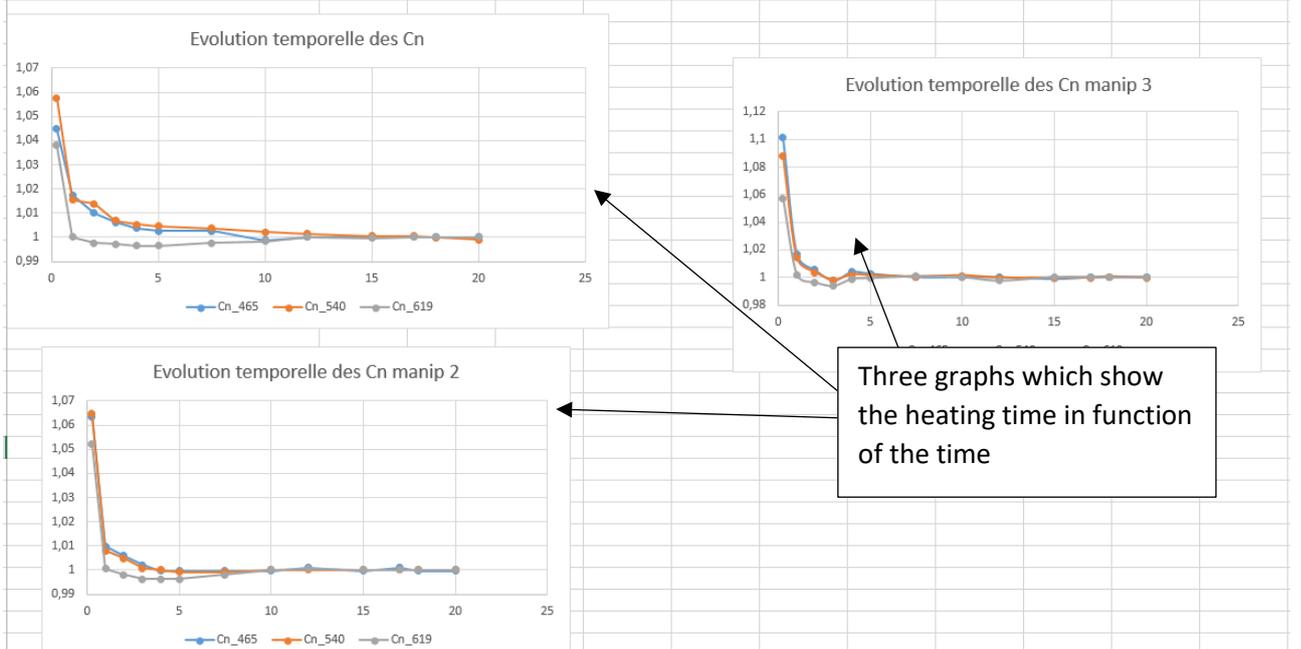
Optical bench used for the calibration



2) Heating time

Before beginning the measures, we have to know the heating time of the lamp. Indeed, we can't do measures if the lamp is not stable. So in order to know it, we'll take measurements just after turning on the lamp. And we found the following results:

Temps après allumage de la lampe (minutes)	Cn_465	Cn_540	Cn_619	Temps après Cn_465	Cn_540	Cn_619	Temps après Cn_465	Cn_540	Cn_619		
0,25	1,04477612	1,05769607	1,03809672	0,25	1,06349703	1,06466877	1,05229994	0,25	1,10110236	1,08774015	1,05726872
1	1,01741294	1,01545081	1,0001555	1	1,00969659	1,00788644	1,00063012	1	1,01669291	1,01406219	1,00188798
2	1,00995025	1,01388617	0,99766755	2	1,00594307	1,00473186	0,99810964	2	1,00535433	1,00376312	0,99622404
3	1,00621891	1,0068453	0,99704556	3	1,00218955	1,00078864	0,99621928	3	0,99779528	0,99821747	0,99370673
4	1,00373134	1,00528066	0,99642357	4	0,99968721	1	0,99621928	4	1,00409449	1,00217865	0,99874135
5	1,00248756	1,00449834	0,99642357	5	0,99968721	0,99921136	0,99621928	5	1,00283465	1,00138641	0,99937067
7,5	1,00248756	1,00371602	0,99766755	7,5	0,99968721	0,99921136	0,99810964	7,5	1,00031496	1,00059418	1,00062933
10	0,99875622	1,00215138	0,99828954	10	0,99968721	1	1	10	1,00031496	1,00138641	1
12	1	1,00136906	1,0001555	12	1,00093838	1	1,00063012	12	1,00031496	0,99980194	0,99748269
15	1	1,00058674	0,99953351	15	0,99968721	1	1	15	0,99905512	0,99980194	1
17	1	1,00058674	1,0001555	17	1,00093838	1	1	17	1,00031496	0,99980194	1
18	1	0,99980442	1,0001555	18	0,99968721	1	1	18	1,00031496	1,00059418	1
20	1	0,9990221	1,0001555	20	0,99968721	1	1	20	1,00031496	0,99980194	1



To determine this time, we place the Calitoo in front of the lamp, and after that, we turn on the lamp and we take some measures at different moments. The aim is to see from what moment the values of the Cn don't vary. For these graphs, we normalize it in order to be able to compare between the different wavelengths.

On these graphs, we can determine the heating time of this lamp. We can see that after 10 minutes the Cn don't vary more than 0.5% (which represent 1Cn for channel blue, 4 for the green and 3 for the red one). After 15 minutes the variation is only about 0.3% so we can say that the heating time of that lamp is about 15 minutes.

3) Flow of the method

Firstly we have to present how we will take our measures. So in order to be as precise as possible, we choose a master and take all the other Calitoo we want to calibrate. After that we do a set of measures (in general 10) with the master, then a set with the first we have to calibrate, the master again and that until the last photometer to calibrate and we finish with a set of measures with the master.

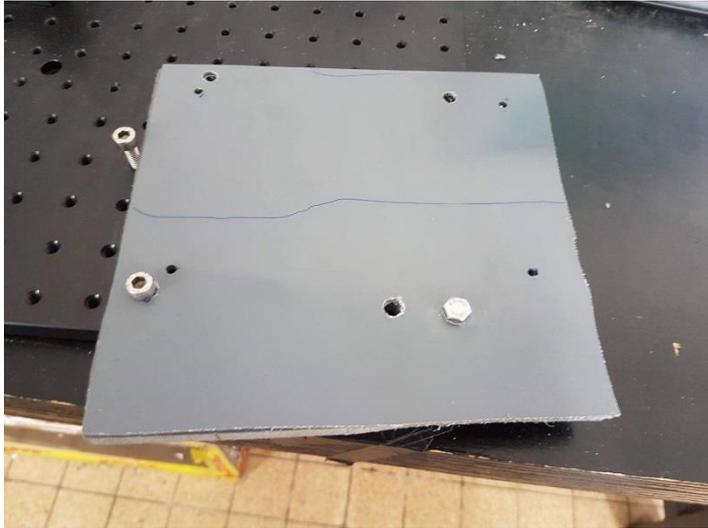
It is also important to know the value of the "black". To do that, we turn off the lamp and we take a measure. We see that the value is 0 and so the influence of other source of light has no impact on the values seen by the Calitoo during the calibration.

4) Problems we have with that method

-Position of the Calitoo

So as to have the most precise measures as possible, we have to put the Calitoo in the better place (where we have the maximum of intensity). So firstly, we found the distance between the Calitoo and the lamp thanks to the Cn which are big enough to analyze but not too close to the lamp to avoid burning. With the sun, the values are close to 2000-2500 if the sky is clear and with the lamp these values are around 700-1600. To make the best measures as possible we have not to be under 1000 if it is possible. This value represents the value of measures on the sun/2. So we decided a place for the Calitoo which is a good compromise. But distance is not the only parameter to take into account. We have to find also the high and the different angles to place the Calitoo.

To be sure of the position of the Calitoo in order to repeat the method, we can put a stop where the Calitoo has to be (for now we use a screw). After that, we can move the directions of the Calitoo to find the maximum but the distance between the lamp and the Calitoo must be the same.



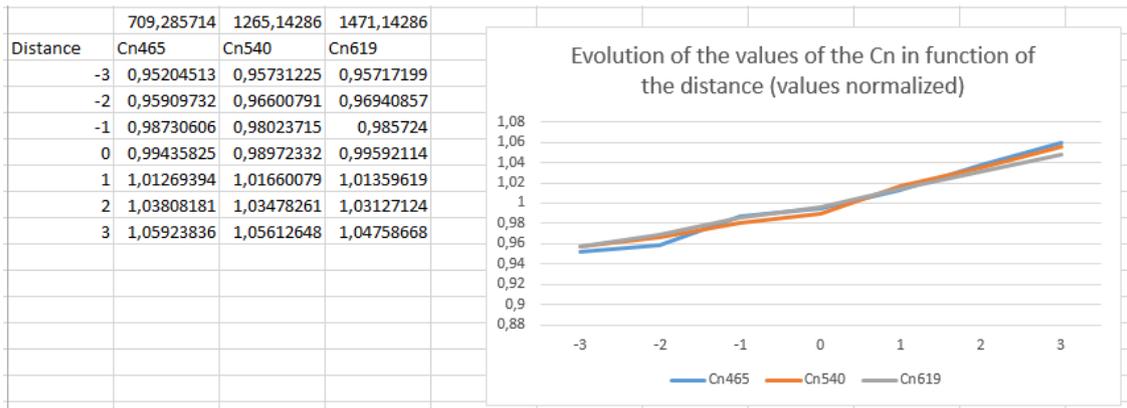
Where we place the Calitoo with the stop in order to fix the distance



We can also notice that the values of the C_n between the three wavelengths are not in the same order with the sun and with the lamp. In fact with the sun the maximum is for the blue channel and the minimum for the red one and it is the opposite for the lamp. We can explain that with the spectrum of the sun and the lamp. The lamp has few power for the channel in UV in contrary to the sun. In fact it is quite impossible to have a lamp which deliver the same wavelengths as the sun because UV is dangerous for our skin and so constructors of the lamps block these UV.

We have here to find a maximum on the C_n of the Calitoo but we noticed that that position is not the same between the different calitoo so we can't find a 'position of calibration'. So we have to find manually that position of maximum for every measure. We can change the different angles of the Calitoo but the distance never changes.

Some errors on the position can affect the values of the C_n like the angles or the high. So here is a recapitulation of the impacts of these parameters:



The evaluation of the impacts of the angles is more difficult to put in place and we can find by hand the maximum possible at every measure. The error we can have due to the position of the Calitoo is about 4 Cn so 0.3%.

-Sensibility of the supply

We have also to be careful with the supply and not to change the value of the tension issued by the supply because the values of the Cn seen by the calitoo change with that tension. And so we can't compare some values if they are not done at the same tension.

-Choice of the reference

In order to do our calibration, we need a photometer reference or master. For the moment we use a calitoo which is supposed to be calibrated but we don't really know the precision of it instead is values are close to those of aeronet. The importance of that photometer is crucial because all the calibration depend on it. If it is possible we can also put another photometer like those in the roof of the laboratory in front of our lamp and do the same job but it is not for the moment.

-Position of all the elements

During a manipulation it is important not to move the different elements on the optical bench to be able to analyze our results.

5) Presentation of the method

Firstly, we have to know if our new method is stable and repeatable. In order to do that, we can take some measures with one Calitoo at different moments. Once we have these data, we can calculate the standard deviation of the Cn. We have done it for some data and find:

CALITOO #1311-0071 Master			
Vo master			
3734			
3322			
2645			
Moyenne	791,6666667	1259	1583,58333
Ecart type	1,775250729	5,04524979	3,47610894
Pourcentage	0,22%	0,40%	0,22%
CALITOO #1311-0062			
3597			
3155			
2560			
Moyenne	793,2	1222,4	1593,2
			Rapports #62/master#71
			1,00193684 0,97092931 1,00607273
Ecart type	0,447213595	0,54772256	0,83666003
Pourcentage	0,06%	0,04%	0,05%
CALITOO #1311-0045			
3808			
3553			
2692			
Moyenne	786,25	1320,5	1584
			Rapports #45/maste#71
			0,99315789 1,04884829 1,00026312
Ecart type	2,5	0,57735027	0,81649658
Pourcentage	0,32%	0,04%	0,05%

Data and constants from the master #71 (15 measures)

Data and constants for the field Calitoo #62 (5 measures)

Data and constants for the field Calitoo #45 (5 measures)

We can see that the standard deviation in percentage is less than 0.5% for these three Calitoo. So the stability and repeatability of this method is not a problem for the following of the process.

With this method, we don't need to do a temporal interpolation because after we get the stability of the lamp all the measures are done in the same conditions.

To validate this calibration method of Calitoo, we have to compare it to some other methods. We know how to calibrate the calitoo thanks to the sun and thanks to the soft of calitoo. If we compare the new Cn0 that we get with the soft and with the new method we have:

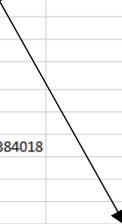
				465nm	540nm	619nm	
CALITOO #1311-0062			old V0	3597	3155	2560	
New V0 lab calibration	11/07/2017			V0 calculated by Calitoo			
3741,23217	3225,42716	2661,06236		3710	3146	2556	
Difference between old and new V0 (lab calibration)				Difference between old and new V0 by calitoo			
144,232168	70,4271644	101,062359		113	-9	-4	
Difference between these two new V0							
31,2321684	79,4271644	105,062359					
CALITOO #1311-0045			old V0	3808	3553	2692	
New V0 lab calibration	11/07/2017			V0 calculated by Calitoo			
3708,45158	3484,27403	2645,69594		3605	3361	2542	
Difference between old and new V0 (lab calibration)				Difference between old and new V0 by calitoo			
-99,5484211	-68,725973	-46,3040573		-203	-192	-150	
Difference between these two new V0							
103,451579	123,274027	103,695943					

The difference between the two different new CnO is important but we have always the one calculated in laboratory higher than the one given by the soft. If we take a look at the signs, we can see that there are always the same and that is a first good point for our method. But these values are not close enough to those of the soft Calitoo. So we have to investigate a little bit more in order to know where the problem is. In a first time, this method can be done with other Calitoo. We don't really know how good these ones are so we will check that with others just calibrated from Izana.

So we did the same experimentation with some other Calitoo taking for master a Calitoo calibrated from Izana. Here are some results:

CALITOO #1703-0236 Master									
Vo master									
3481									
3477									
2580									
Moyenne	706,2	1267,8	1467,9						
Ecart type	2,043961296	4,39191176	1,59513148						
Pourcentage	0,29%	0,35%	0,11%						
CALITOO #1703-0237									
3298									
3121									
2528									
Moyenne	668,3	1127,1	1430,2	Rapports #237/master#236			New V0		
				0,94633248	0,88902035	0,97431705	3294,18338	3091,12376	2513,73799
Ecart type	3,713339318	6,52261025	7,85705628				Difference between old and new V0		
Pourcentage	0,56%	0,58%	0,55%				-3,81662419	-29,8762423	-14,2620069
							-0,12%	-0,97%	-0,57%
CALITOO #1703-0238									
3361									
3155									
2568									
Moyenne	671,1	1126,7	1429,5	Rapports #238/master#236			New V0		
				0,95029737	0,88870484	0,97384018	3307,98513	3090,02674	2512,50766
Ecart type	2,330951165	4,71522357	3,13581462				Difference between old and new V0		
Pourcentage	0,35%	0,42%	0,22%				-53,0148683	-64,9732608	-55,492336
							-1,60%	-2,10%	-2,21%

Percentage of difference



Data of the three Calitoo just calibrated from Izana 19/07/2017

The data for a first manipulation show that the standard deviation is correct for both of the field Calitoo. The percentage of difference between the old and the new V0 is also good (less than 1% of maximum for the #237). We can compare these measures with the same but in other days in order to know if it is repeatable.

The first results are here:

CALITOO #1703-0238 Master								
Vo master								
3361								
3155								
2568								
Moyenne	671,1	1126,7	1429,5					
Ecart type	2,330951165	4,71522357	3,13581462					
Pourcentage	0,35%	0,42%	0,22%					
CALITOO #1504-0188								
3626								
3182								
2552								
Moyenne	705,6	1112,3	1417,4	Rapports #188/master#236				CALITOO #1504-0188
				1,05140814	0,98721931	0,9915355	New V0 lab calibration	
							3533,78274	3114,67693
							2546,26317	
Ecart type	5,719362824	8,1520277	3,20416396				Difference between old and new V0	
Pourcentage	0,81%	0,73%	0,23%				-92,2172553	-67,3230674
							-5,73683106	
							-2,61%	-2,16%
							-0,23%	
CALITOO #1504-0184								
3788								
3369								
2624								
Moyenne	745,8333333	1188,08333	1461	Rapports #184/master#236				CALITOO #1504-0184
				1,11135946	1,05448064	1,02203568	New V0 lab calibration	
							3735,27914	3326,88641
							2624,58762	
Ecart type	2,208797836	3,20392751	8,29019136				Difference between old and new V0	
Pourcentage	0,30%	0,27%	0,57%				-52,7208563	-42,1135913
							0,58761805	
							-1,41%	-1,27%
							0,02%	

Percentage of difference

Data taking as master one of the Calitoo from Izana and as fields some of Calitoo never used
20/07/2017

We can see here that the standard deviation is always less than 1% for every Calitoo. The difference between old and new CnO in percentage is not very high. We are under 3% so it could be acceptable for these Calitoo.

To finish the validation of the method, we have to know if the method is also valuable for some Calitoo that have been used. Comparing with the new Cn0 calculated by the soft of Calitoo, we can know if we are close or not from these values and so know if we can apply this method to every Calitoo.

So we have done the same protocol with the same master and with field instrument the #71 and the #45.

CALITOO #1703-0236			Master			
Vo master						
3481						
3477						
2580						
Moyenne						
	701,6666667	1256,86667	1458,6			
Ecart type	1,632993162	3,37779866	6,17367684			
Pourcentage	0,23%	0,27%	0,42%			
CALITOO #1311-0071						
3734						
3322						
2645						
Moyenne				Rapports #71/master#2336		
	763,6363636	1206,18182	1525,63636	1,08831786	0,95967365	1,04595939
Ecart type	4,674884539	7,30504185	4,67488454			
Pourcentage	0,61%	0,61%	0,31%			
CALITOO #1311-0062						
3597						
3155						
2560						
Moyenne				Rapports #62/master#236		
	775,6	1186,4	1551	1,10536817	0,94393465	1,06334842
Ecart type	1,429840706	3,40587727	3,19722102			
Pourcentage	0,18%	0,29%	0,21%			

Data taking as master a Calitoo from Izana and as fields two of Calitoo used previously 20/07/2017

So the standard deviation is under 0.6% for these two Calitoo as we can see here.

				465nm	540nm	619nm	
CALITOO #1311-0071			old V0	3734	3322	2645	
New V0				V0 calculated by Calitoo			
3788,43446	3336,78527	2698,57522		3796	3348	2686	
Difference between old and new V0				Difference between old and new V0 by calitoo			
54,4344634	14,7852717	53,5752216		62	26	41	
1,44%	0,44%	1,99%					
Difference between these two new V0							
-7,5655366	-11,2147283	12,5752216					
-0,20%	-0,34%	0,48%					
CALITOO #1311-0062			old V0	3597	3155	2560	
New V0				V0 calculated by Calitoo			
3847,7866	3282,06079	2743,43891		3790	3241	2713	
Difference between old and new V0				Difference between old and new V0 by calitoo			
250,786603	127,060786	183,438914		193	86	153	
6,52%	3,87%	6,69%					
Difference between these two new V0							
57,7866033	41,0607861	30,438914					
1,61%	1,30%	1,19%					

Data taking as master a Calitoo from Izana and as fields two of Calitoo used previously
20/07/2017

Here we can't study the difference between the old and the new Cn0 because these Calitoo are not calibrated. But thanks to the soft of Calitoo we can calculate a new Cn0 and so we can compare the two new Cn0 in order to know if the results of our method are similar to those of Calitoo.

We see a difference of less than 0.5% for the first Calitoo and less than 2% for the second one. So the values calculated by Calitoo and by our method are very close. In fact if we do the calibration with Calitoo or with our method we find more or less the same values of Cn0. This is good for the validation of the method.

Conclusion

To conclude we can say that the results we have seen previously show the future possibilities of that method. For the moment our results seem to validate the method because we have seen the repeatability and the stability of the method in addition to the previous results. But we don't have enough measures in order to know if we can do that for every Calitoo. Some parameters like the angles of the Calitoo or the precision of the lamp are not taken in count and if we want to validate all the method we have to understand what is the impact of these parameters.

We can recapitulate now the different errors which can affect our measures. Here is a list of these errors:

- stability of the lamp (1-2 Cn 0.1%)
- stability of the alimentation (1 Cn 0.1%)
- position of the Calitoo: angles, high,... (4 Cn 0.4%)
- precision of the Calitoo (0.02 on the AOD)

So we have an error of: $\sqrt{2^2 + 1^2 + 4^2 + 2^2} = \sqrt{25} = 5$ on the Cn which represent less than 0.5% of error for each channel. This possible error on the Cn represent an error on the AOD of maximum 0.015.

Some atmospheric corrections (Rayleigh and Ozone) appear in the calculation of the AOD. These corrections are constant for each site but can introduce an error when the intercalibration or the Langley calibration is done. Doing the calibration in laboratory these correction don't come in count. This is one of the difference between the methods and advantage for the method in laboratory.

We have also to take in count the master because until now, we use a Calitoo as master but it will be better to use another photometer. A changement of the master could be considered in the future and two options can be considered: taking one of the CIMEL photometer or taking a MICROTOPS photometer.